

# 237(5): The Cotes and Poincaré Spirals and Frechet

## Analysis of the Kinematics of Orbits

It has been shown that the Coriolis acceleration vanishes in all orbits, so the force law of all orbits is:

$$\underline{F}(r) = m \left( \frac{d^2 r}{dt^2} \underline{e}_r + \underline{\omega} \times (\underline{\omega} \times \underline{r}) \right) = -\frac{L^2}{mr^3} \left( \frac{d^2}{dt^2} \left( \frac{1}{r} \right) + \frac{1}{r} \right) \underline{e}_r \quad - (1)$$

where

$$\underline{L} = \underline{r} \times \underline{p}, \quad - (2)$$

$$L = |\underline{L}| = mr^2 \omega. \quad - (3)$$

The magnitude of the spin convention is:

$$\omega = |\underline{\omega}| = \frac{d\theta}{dt}. \quad - (4)$$

So the force depends entirely on the spin convention and the rotation of plane polar axes.

The two components of the force are:

$$1) \quad \underline{F}_1 = \frac{d^2 r}{dt^2} \underline{e}_r = -\frac{L^2}{mr^3} \frac{d^2}{dt^2} \left( \frac{1}{r} \right) \underline{e}_r \quad - (5)$$

and:

$$2) \quad \underline{F}_2 = \underline{\omega} \times (\underline{\omega} \times \underline{r}) = -\frac{L^2}{mr^3} \underline{e}_r \quad - (6)$$

Note carefully that the inverse square law of

2) Hooke and Newton is the sum of  $\underline{F}_1$  and  $\underline{F}_2$ . The second type of force law (b) does not depend on the  $r(\theta)$  function, it is the same for all orbits in a plane and is known as the centrifugal force. The latter is due to the sign convention  $\omega$  and to the rotation of axes. The first type of force  $\underline{F}_1$  is the inertial force and does not depend on rotation of axes.

The centrifugal force does not exist in a static frame in which the axes are constant. Therefore the Hooke/Newton inverse square law applied to the attraction between  $m$  and  $M$  is:

$$\underline{F} = \underline{F}_1 = -\frac{mMG}{r^2} \underline{e}_r \quad (7)$$

However, for an elliptical orbit:

$$\frac{1}{r} = \frac{1}{d} (1 + \epsilon \cos \theta) \quad (8)$$

and

$$\frac{d^2}{d\theta^2} \left( \frac{1}{r} \right) = -\frac{\epsilon}{d} \cos \theta \quad (9)$$

so

$$\begin{aligned} \underline{F}_1 &= \frac{L^2}{mr^3} \frac{\epsilon \cos \theta}{d} \underline{e}_r \quad (10) \\ &= \frac{L^2}{mr^3} \left( \frac{1}{r} - \frac{1}{d} \right) \underline{e}_r \end{aligned}$$

3) Therefore when the axes start moving the force  $F_1$  becomes different. The total force for an orbit of type (8) is:

$$\begin{aligned}\underline{F} &= \underline{F}_1 + \underline{F}_2 \\ &= \frac{L^2}{mr^3} \left( \frac{1}{r} - \frac{1}{d} \right) \underline{e}_r - \frac{L^2}{mr^3} \underline{e}_r \\ &= - \frac{L^2}{mr^3 d} \underline{e}_r. \quad - (11)\end{aligned}$$

$$= m \frac{d^2 r}{dt^2} \underline{e}_r + m \underline{\omega} \times (\underline{\omega} \times \underline{r})$$

The origin of the appellation "universal gravitation" is that eq. (11) has a similar structure to eq. (7). However eq. (7) contains no centrifugal acceleration.

In order for eq. (11) to become (7):

$$d = \frac{L^2}{m^2 M G} \quad - (12)$$

Universal gravitation exists if and only if  
 $d$  is assumed to be eq. (12). This is true only  
for an elliptical orbit. The observed orbits of  
astronomy are not elliptical.



For example the (otes spirals are:

$$\frac{1}{r} = A \cos(k\theta + \epsilon), \quad - (13)$$

$$\frac{1}{r} = A \cosh(k\theta + \epsilon), \quad - (14)$$

$$\frac{1}{r} = A\theta + \epsilon \quad - (15)$$

and the Poincaré spirals are:

$$\frac{1}{r} = \frac{1}{a} \cosh(n\theta), \quad - (16)$$

$$\frac{1}{r} = \frac{1}{a} \sinh(n\theta). \quad - (17)$$

From eq. (13):

$$\underline{F}(r) = -\frac{L^2}{mr^3} (1 - k^2) \underline{e}_r \quad - (18)$$

so:

$$\underline{F}_1 = \frac{k^2 L^2}{mr^3}, \quad \underline{F}_2 = -\frac{L^2}{mr^3} \quad - (19)$$

and

$$\underline{F} = \left( \frac{k^2 L^2}{mr^3} - \frac{L^2}{mr^3} \right) \underline{e}_r \quad - (20)$$

$$= \frac{k^2 L^2}{mr^3} \underline{e}_r + \frac{m}{L} \underline{\omega} \times (\underline{\omega} \times \underline{r})$$

The "universal" law of gravitation no longer holds, but the centrifugal force is the same.

For the Poincaré spiral (16):

5)

$$\underline{F}(r) = -\frac{L^2}{mr^3} (1 + k^2) \underline{e}_r \quad - (21)$$

$$\underline{F}_1 = -\frac{k^2 L^2}{mr^3} \underline{e}_r \quad - (22)$$

$$\underline{F}_2 = m \underline{\omega} \times (\underline{\omega} \times \underline{r}) = -\frac{L^2}{mr^3} \underline{e}_r \quad - (23)$$

The Cotes spiral (14) is the Poisson spiral (16) if:

$$A = \frac{1}{a}, \quad k = n, \quad \epsilon = 0, \quad - (24)$$

and the Cotes spiral (13) written in the form:

$$f(\theta) = \frac{\epsilon}{d} \cos(x\theta) \quad - (25)$$

becomes the precessing ellipse if:

$$\frac{1}{r} = \frac{1}{d} + f(\theta). \quad - (26)$$

So the precessing ellipse is the Poisson spiral (25) added to  $1/d$ .

For the precessing ellipse:

$$\frac{d^2}{d\theta^2} \left( \frac{1}{r} \right) = -x^2 \frac{\epsilon}{d} \cos(x\theta) \quad - (27)$$

so from eq. (5):

$$\begin{aligned}
 \underline{F}_1 &= \frac{L^2 x^2}{m d r^2} \cos(x\theta) \underline{e}_r \\
 &= \frac{L^2 x^2}{m r^2 d} \left( \frac{d}{r} - 1 \right) \underline{e}_r \\
 &= \frac{L^2 x^2}{m r^2} \left( \frac{1}{r} - \frac{1}{d} \right) \underline{e}_r \quad - (28)
 \end{aligned}$$

so the complete force law is:

$$\begin{aligned}
 \underline{F} &= \left( x^2 \frac{L^2}{m r^2} \left( \frac{1}{r} - \frac{1}{d} \right) - \frac{L^2}{m r^3} \right) \underline{e}_r \\
 &= \left( (x^2 - 1) \frac{L^2}{m r^3} - \frac{x^2 L^2}{m r^2 d} \right) \underline{e}_r \quad - (29)
 \end{aligned}$$

which is the sum of inverse square and cube terms.

Each orbit has its different force law and

there is no universal gravitation.

However, the centrifugal force:

$$\underline{F} = m \underline{\omega} \times (\underline{\omega} \times \underline{r}) = - \frac{L^2}{m r^3} \underline{e}_r \quad - (30)$$

is the same for all planar orbits. The force  $\underline{F}$  is directed towards the origin in the observer frame of reference. In the solar system the origin is at  $M$ , the mass of



7) The sun. The force is directed inward because it is the force needed to counterbalance the outward directed centrifugal force  $-\frac{L^2}{mr^3} \underline{e}_r$  of the planet:

$$\underline{F}(\text{inward}) = -m\omega \times (\omega \times \underline{r}) = \frac{L^2}{mr^3} \underline{e}_r - (31)$$

The force  $\underline{F}(\text{outward})$  is the every day force experienced in any rotational motion. For example in the hammer in athletics an inward force:

$$\underline{F}(\text{inward}) = m\omega \times (\omega \times \underline{r}) = -\frac{L^2}{mr^3} \underline{e}_r - (32)$$

is exerted on the hammer. When it is released, it accelerates away from the athlete. A planet in any planar orbit always tries to move in a straight line tangential to the orbit, but is prevented from doing so by the inward force (32) exerted by the sun.

All planar orbits can be analyzed by the

Frenet formulae:

$$\underline{T} = \frac{d\underline{r}}{dt} - (33)$$

$$\frac{d\underline{T}}{dt} = \frac{d^2 \underline{r}}{dt^2} = \frac{1}{\rho} \underline{N} = \kappa \underline{N} - (34)$$

$$\frac{d\underline{B}}{dt} = -\tau \underline{N} - (35)$$

$$\frac{d\mathbf{N}}{dt} = -\frac{1}{\rho} \mathbf{T} + \tau \mathbf{B} \quad - (36)$$

Here  $\mathbf{T}$  is the tangent vector,  $\mathbf{N}$  is the normal vector,  $\kappa$  is the curvature and  $\tau$  is the Frenet torsion. Therefore

the Frenet curvature is:

$$\kappa = \frac{|\dot{\mathbf{r}} \times \ddot{\mathbf{r}}|}{|\dot{\mathbf{r}}|^3} \quad - (37)$$

and the Frenet torsion is:

$$\tau = \frac{\dot{\mathbf{r}} \cdot \ddot{\mathbf{r}} \times \ddot{\mathbf{r}}}{|\dot{\mathbf{r}} \times \ddot{\mathbf{r}}|^2} \quad - (38)$$

The binormal unit vector of Frenet is:

$$\mathbf{B} = \mathbf{T} \times \mathbf{N} \quad - (39)$$

For the hyperbolic spiral orbit:

$$\mathbf{r} = -\frac{r_0}{\theta} \quad - (40)$$

The velocity is:

$$\mathbf{v} = \mathbf{T} = \frac{L}{m} \left( \frac{1}{r_0} \mathbf{e}_r + \frac{1}{r} \mathbf{e}_\theta \right) \quad - (41)$$

and the acceleration is:

$$\mathbf{a} = \frac{1}{\rho} \mathbf{N} = -\frac{L^2}{m^2 r^3} \mathbf{e}_r \quad - (42)$$



9) Therefore:

$$\begin{aligned}\underline{v} \times \underline{a} &= -\frac{L}{mr} \underline{e}_\theta \times \left( \frac{L^2}{m^2 r^3} \right) \underline{e}_r \\ &= -\frac{L^3}{mr^4} \underline{e}_\theta \times \underline{e}_r \\ &= \frac{1}{r} \left( \frac{L}{mr} \right)^3 \underline{k} \quad - (43)\end{aligned}$$

Using

$$L = mr^2 \omega \quad - (44)$$

then

$$\boxed{\underline{v} \times \underline{a} = r^2 \omega^3 \underline{k}} \quad - (45)$$

The Frenet curvature of the hyperbolic spiral is:

$$\boxed{\kappa = \frac{|\underline{v} \times \underline{a}|}{|\underline{v}|^3} = \frac{1}{r} \left( \frac{\omega r}{v} \right)^3} \quad - (46)$$

The Biconal is:

$$\underline{B} = \underline{T} \times \underline{N} = \rho \underline{v} \times \underline{a} = \frac{\rho}{r} \left( \frac{\omega r}{v} \right)^3 \underline{k} \quad - (47)$$

The normal is:

$$\underline{N} = \rho \underline{a} = -\rho \frac{L^2}{m^2 r^3} \underline{e}_r \quad - (48)$$

By definition:

$$\underline{T} = \frac{d\underline{r}}{dt} = \frac{dx}{dt} \underline{i} + \frac{dy}{dt} \underline{j} \quad - (49)$$

10) and

$$\frac{d\mathbf{T}}{dt} = \kappa \mathbf{N} \quad - (50)$$

so

$$N_x = \frac{1}{\kappa} \frac{dT_x}{dt} \quad - (51)$$

$$N_y = \frac{1}{\kappa} \frac{dT_y}{dt} \quad - (52)$$

$$N_z = 0 \quad - (53)$$

Therefore:

$$\begin{aligned} \mathbf{B} = \mathbf{T} \times \mathbf{N} &= (T_y N_z - T_z N_y) \mathbf{i} \\ &+ (T_z N_x - T_x N_z) \mathbf{j} \\ &+ (T_x N_y - T_y N_x) \mathbf{k} \quad - (54) \\ &= (T_x N_y - T_y N_x) \mathbf{k} \end{aligned}$$

Finally:

$$\frac{d\mathbf{B}}{dt} = -\tau \mathbf{N} \quad - (55)$$

$$\text{so } \left( 0, 0, \frac{dB_z}{dt} \right) = -\tau \left( \frac{dN_x}{dt}, \frac{dN_y}{dt}, 0 \right)$$

so for all planar cs. it's @ Frenet torsion  
vanishes.

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