

2.10(3): Transition to Planar Orbit

The basic equations of the 3-D orbit are:

$$\mathcal{L} = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\beta}^2) + V(r) \quad (1)$$

where $V(r)$ is any central potential of attraction. For the inverse square force:

$$V(r) = -\frac{k}{r} \quad (2)$$

Here:

$$\dot{\beta}^2 = \dot{\theta}^2 + \dot{\phi}^2 \sin^2 \theta \quad (3)$$

The force of attraction is:

$$F(r) = -\frac{\partial V(r)}{\partial r} \quad (4)$$

The Euler Lagrange equations are:

$$\frac{\partial \mathcal{L}}{\partial r} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{r}} \quad (5)$$

$$\frac{\partial \mathcal{L}}{\partial \beta} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\beta}} \quad (6)$$

$$\frac{\partial \mathcal{L}}{\partial \theta} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\theta}} \quad (7)$$

$$\frac{\partial \mathcal{L}}{\partial \phi} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\phi}} \quad (8)$$

2) Eq. (5) gives:

$$m\ddot{r} = mr\dot{\beta}^2 - \frac{\partial V(r)}{\partial r} \quad - (9)$$

which is the Leibniz equation.

Eqs. (6) and (7) give:

$$\frac{d\beta}{dt} = \frac{L}{mr^2} \quad - (10)$$

and

$$\frac{d\theta}{dt} = \frac{L_\theta}{mr^2} \quad - (11)$$

which imply:

$$\frac{d\phi}{dt} = \frac{L_\phi}{mr^2 \sin \theta} \quad - (12)$$

where

$$L^2 = L_\theta^2 + L_\phi^2 \quad - (13)$$

The Euler Lagrange equation (8) gives:

$$\ddot{\phi} \sin \theta + 2\dot{\theta} \dot{\phi} \cos \theta = 0 \quad - (14)$$

From eqs. (10) and (11):

$$\frac{d\beta}{d\theta} = \frac{L}{L_\theta} \quad - (15)$$

From eqs. (10) and (12):

$$\frac{d\beta}{d\phi} = \frac{L}{L_\phi} \sin \theta \quad - (16)$$

3) If constants of integration are assumed to be zero then:

$$\beta = \frac{L}{L_\theta} \dot{\theta} = \frac{L}{L_\phi} \dot{\phi} \sin \theta - (17)$$

as in UFT 269. However, from eq. (17):

$$\dot{\beta} = \frac{L}{L_\theta} \ddot{\theta} = \frac{L}{L_\phi} (\dot{\phi} \sin \theta + \phi \dot{\theta} \cos \theta) - (18)$$

so

$$\ddot{\theta} = \frac{L_\theta}{L_\phi} (\dot{\phi} \sin \theta + \phi \dot{\theta} \cos \theta) - (19)$$

From (11), (12) and (19)

$$L_\theta = \frac{L_\theta}{L_\phi} (L_\phi + \phi \cos \theta L_\theta) - (20)$$

so

$$\phi \cos \theta = 0 - (21)$$

and

$$\theta = \pi/2 - (22)$$

So the solution (17) leads back to a planar orbit. Eq. (17) is therefore the reason for planar orbits.

In order to obtain non planar orbits

4) constants of integration must be used, in the simplest case:

$$\beta = \frac{L}{L_0} \theta + A(t) - (23)$$

and

$$\beta = \frac{L}{L_\phi} \phi \sin \theta - (24)$$

where $A(t)$ is time dependent but not θ dependent. Then eq. (18) is no longer true and θ is no longer fixed at $\pi/2$.

Three Dimensional Orbits in General

In general choose the solution (24), so for an inverse square law:

$$r = \frac{a}{1 + e \cos\left(\frac{L}{L_\phi} \sin \theta \phi\right)} - (25)$$

and previous results are all valid.

For the general case such as Φ_{exp} in 3-D galaxies the 3-D Binet equation is:

$$F(r) = -\frac{\partial V(r)}{\partial r} = -\frac{L^2}{mr^3} \left(\frac{d^2}{d\beta^2} \left(\frac{1}{r} \right) + \frac{1}{r} \right) - (26)$$

) which:

$$\beta = \frac{L}{L_\phi} \phi \sin \theta \quad - (27)$$

and

$$m \ddot{r} = - \frac{L^2}{m r^3} \frac{d^2}{d\beta^2} \left(\frac{1}{r} \right) \quad - (28)$$

For the inverse square law of attraction:

$$F(r) = - \frac{k}{r^2} \quad - (29)$$

The orbit is:

$$r = \frac{\alpha}{1 + \epsilon \cos \beta} \quad - (30)$$

which is a precessing ellipse:

$$r = \frac{\alpha}{1 + \epsilon \cos \left(\left(\frac{L}{L_\phi} \sin \theta \right) \phi \right)} \quad - (31)$$

Under the condition (17), i.e.:

$$\theta = \frac{L_\theta}{L_\phi} \phi \sin \theta \quad - (32)$$

eq (31) reduces to the planar ellipse:

$$r = \frac{\alpha}{1 + \epsilon \cos \phi} \quad - (33)$$

for all L_θ and L_ϕ .

1) For other types of force of attraction, the 3-D Binet equation (26) produces different three dimensional orbits. These orbits are observed in three dimensional galaxies. They depend on the function:

$$f(r, \phi, \theta) = \frac{d^2}{d\beta^2} \left(\frac{1}{r} \right) - (27)$$

where

$$\beta = \frac{L}{L\phi} \phi \sin\theta - (28)$$

For a given choice of orbit the force law can be determined. Since β is a function of both ϕ and θ , this is a problem in the function of two variables, so the chain rule of differential calculus is needed.
