

403(10) : Development of Note 403(6)

In the near circular approximation, the apsidal method gives an orbital precession of:

$$\Delta\phi = \frac{1}{3} \frac{\langle \underline{sr} \cdot \underline{sr} \rangle}{r^2} - \frac{r^2}{2} \frac{d\omega}{dr}, \quad - (1)$$

which is eq. (41) of Note 403(6). In eq. (1):

$$\omega = \frac{2}{3} \frac{\langle \underline{sr} \cdot \underline{sr} \rangle}{r^3} \quad - (2)$$

so

$$\frac{d\omega}{dr} = -2 \frac{\langle \underline{sr} \cdot \underline{sr} \rangle}{r^4} + \frac{2}{3} \frac{1}{r^3} \frac{d}{dr} \langle \underline{sr} \cdot \underline{sr} \rangle \quad - (3)$$

So:

$$\Delta\phi = \frac{\langle \underline{sr} \cdot \underline{sr} \rangle}{r^2} - \frac{1}{3r} \frac{d}{dr} \langle \underline{sr} \cdot \underline{sr} \rangle$$

- (4)

At the perihelion:

$$r = r_{\min} = a(1-\epsilon) = \frac{d}{1+\epsilon} \quad - (5)$$

where a is the semimajor axis, d is the half right distance and ϵ the eccentricity. In the limit of an exactly circular orbit:

$$\epsilon \rightarrow 0 \quad - (6)$$

so

$$r_{\min} = a = d \quad - (7)$$

If it is assumed that for an exactly circular orbit:

$$\Delta\phi = 0 \quad - (8)$$

2) then eq. (4) becomes:

$$\frac{1}{3a} \frac{d}{dr} \langle \underline{S}_r \cdot \underline{S}_r \rangle = \frac{\langle \underline{S}_r \cdot \underline{S}_r \rangle}{a^2} - (9)$$

so.

$$\frac{d}{dr} \langle \underline{S}_r \cdot \underline{S}_r \rangle = \frac{3}{a} \langle \underline{S}_r \cdot \underline{S}_r \rangle - (10)$$

and

$$\Delta \phi = 0 - (11)$$

Eq. (11) follows from eqs. (4) and (10).

If it is assumed that eq. (10) is approximately true in a near circular orbit, then eq. (4) becomes:

$$\Delta \phi \sim \frac{\langle \underline{S}_r \cdot \underline{S}_r \rangle}{r^2} - \frac{1}{ar} \langle \underline{S}_r \cdot \underline{S}_r \rangle.$$

$$= \frac{\langle \underline{S}_r \cdot \underline{S}_r \rangle}{r} \left(\frac{1}{r} - \frac{1}{a} \right) - (12)$$

The exact result is eq. (4), and can be compared with eq. (2) of Note 403(9):

$$\Delta \phi = \frac{2}{3} d^{1/2} \langle \underline{S}_r \cdot \underline{S}_r \rangle \left/ \left(\frac{u^2}{-du^2 + 2u + \frac{1}{a}} \right)^{3/2} \log du \right. - (13)$$

which has been integrated numerically by (10) and (13), the quantity Eckardt. Comparing eqs. (4) and (13), the quantity

$\frac{d}{dr} \langle \underline{S}_r \cdot \underline{S}_r \rangle$ can be found from:

$$3) \frac{\langle \underline{s}_r \cdot \underline{s}_r \rangle}{r^3} - \frac{1}{3r} \frac{d}{dr} \langle \underline{s}_r \cdot \underline{s}_r \rangle$$

$$= \frac{2}{3} d^{1/2} \langle \underline{s}_r \cdot \underline{s}_r \rangle \int \left(\frac{u^2}{-du^2 + 2u + \frac{1}{a}} \right)^{3/2} \log_e u \, du$$

Eq. (14) is a comparison of precessions using analytical approximations in the near circular limit. — (14)

The orbit is:

$$\phi = - \int \left(\frac{d}{2\omega_r \log_e u - du^2 + 2u + \frac{1}{a}} \right)^{1/2} du$$

— (15)

In the limit:

$$\omega_r \rightarrow 0 \quad - (16)$$

The orbit (15) is the circ section:

$$r = \frac{d}{1 + \epsilon \cos \phi} \quad - (17)$$

In eq. (15): $\omega_r = \frac{2}{3} \frac{\langle \underline{s}_r \cdot \underline{s}_r \rangle}{r^3} \quad - (18)$

so the well known circ section is perturbed by the isotropically averaged vacuum fluctuation $\langle \underline{s}_r \cdot \underline{s}_r \rangle$.
 the vacuum perturbation produce precession given by
 eq. (14) in the near circular limit.