

ss(6): Relation between "Second Bianchi Identity" and
Einstein Field Equations

The derivation by Einstein of R, field equation in 1915 started from "Second Bianchi identity":

$$D \wedge R^a_b = 0 \quad - (1)$$

which in Riemann notation is:

$$D_\lambda R_{\rho\sigma\mu\nu} + D_\rho R_{\sigma\lambda\mu\nu} + D_\sigma R_{\lambda\rho\mu\nu} = 0 \quad - (2)$$

where: $R_{\rho\sigma\mu\nu} = g_{\rho\lambda} R^\lambda_{\sigma\mu\nu} \quad - (3)$

and $g_{\mu\nu} = g_{\nu\mu} \quad - (4)$

It is now known that these equations are true if and

only if: $T^a = 0, \quad - (5)$

where $\Gamma^k_{\mu\nu} = \Gamma^k_{\nu\mu} \quad - (6)$

At the time (1915), the Cartan torsion (1922) was not known, and Einstein accepted eq. (5) as an identity. In fact it is not an identity, it depends on eq. (5). In 1922 in a later work, Cartan showed that in the presence of torsion:

$$D \wedge T^a = R^a_b \wedge v^b \quad - (7)$$

$$T^a = D \wedge v^a \quad - (8)$$

$$R^a_b = D \wedge \omega^a_b \quad - (9)$$

2) In general:

$$D \wedge T^a = R^a{}_b \wedge \omega^b \neq 0 \quad (10)$$

in the presence of torsion. Therefore it general:

$$D \wedge R^a{}_b := D \wedge (D \wedge \omega^a{}_b) \neq 0 \quad (11)$$

and

$$D \wedge (R^a{}_b \wedge \omega^b) := D \wedge (D \wedge T^a) \neq 0 \quad (12)$$

Cosmology and general relativity should therefore be developed from eqs. (11) and (12), not from eq. (1).

(1). For example, spiral galaxies are due to torsion, and not "dark matter". The latter does not exist in natural philosophy, and is akin to "phlogiston".

The actual field equation of Einstein (1915) is derived from the double index contraction:

$$g^{\nu\sigma} g^{\mu\lambda} (D_\lambda R_{\rho\sigma\mu\nu} + D_\rho R_{\sigma\lambda\mu\nu} + D_\sigma R_{\lambda\rho\mu\nu}) = 0$$

The Einstein field equation is true only in the absence of torsion. (13)

In the presence of torsion the Noether Theorem is also extended.

) The following definitions (see Carroll) are then used in eq. (13):

$$D^\mu R_{\rho\mu} := - (g^{\mu\lambda} D_\lambda) (g^{\nu\sigma} R_{\rho\sigma\mu\nu}) \quad - (14)$$

$$D^\nu R_{\rho\nu} := - (g^{\nu\sigma} D_\sigma) (g^{\mu\lambda} R_{\sigma\lambda\rho\mu}) \quad - (15)$$

$$D_\rho R := D_\rho (g^{\nu\sigma} g^{\mu\lambda} R_{\sigma\lambda\rho\mu}) \quad - (16)$$

The convention used by Einstein is as follows. If indices are in the same order as the metric and is a tensor multiplied by the metric, then the resulting sign is positive. If indices are in the opposite order as the tensor to the index order in the metric, then the sign is negative. Thus:

$$D^\mu R_{\rho\mu} - D_\rho R + D^\nu R_{\rho\nu} = 0 \quad - (17)$$

i.e.
$$D^\mu R_{\rho\mu} - \frac{1}{2} D_\rho R = 0 \quad - (18)$$

or
$$R_{\rho\nu} = \frac{1}{4} R g_{\rho\nu}. \quad - (19)$$

Finally we:
$$D_\rho = g_{\rho\mu} D^\mu \quad - (20)$$

to obtain:
$$\boxed{D^\mu \left(R_{\rho\mu} - \frac{1}{2} R g_{\rho\mu} \right) = 0} \quad - (21)$$

) This is a re-emergence of eq. (1). Eq (21) is true if and only if \mathbb{E} torsion is zero. The usual nomenclature is as follows. The Einstein tensor

is:

$$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} \quad - (22)$$

where $R_{\mu\nu}$ is \mathbb{E} Ricci tensor, R is \mathbb{E} scalar curvature, and $g_{\mu\nu}$ is \mathbb{E} metric. Thus:

$$\boxed{D^\mu G_{\mu\nu} = 0} \quad - (23)$$

The conventional (torsion free) Noether tensor is:

$$\boxed{D^\mu T_{\mu\nu} = 0} \quad - (24)$$

where $T_{\mu\nu}$ is \mathbb{E} canonical energy-momentum tensor.

Due to reflect of torsion:

$$G_{\mu\nu} = G_{\nu\mu} \quad - (25)$$

$$T_{\mu\nu} = T_{\nu\mu} \quad - (26)$$

To star \mathbb{E} famous field equation, Einstein made eq. (23) proportional to eq. (24).

$$D^\mu g_{\mu\nu} = k D^\mu T_{\mu\nu} = 0 \quad - (27)$$

and so:

$$\boxed{g_{\mu\nu} = k T_{\mu\nu}} \quad - (28)$$

where k is Einstein constant.

Modern cosmology is all based on eq. (28)
It completely reflects Kasica.

Eq. (28) can equally well be written as:

$$D \wedge R^a{}_b = k D \wedge T^a{}_b = 0 \quad - (29)$$

where: $R^a{}_b := R^a{}_{b\mu\nu} \quad - (30)$

is Riemann curvature form and:

$$T^a{}_b = T^a{}_{b\mu\nu} \quad - (31)$$

is Noether form. In presence of Kasica:

$$D \wedge R^a{}_b = k D \wedge T^a{}_b \neq 0 \quad - (32)$$

and

$$\boxed{R^a{}_{b\mu\nu} = k T^a{}_{b\mu\nu}} \quad - (33)$$

In general:

$$R^a{}_{b\mu\nu} \neq R^b{}_{a\mu\nu} \quad - (34)$$

$$T^a{}_{b\mu\nu} \neq T^b{}_{a\mu\nu} \quad - (35)$$

b) The usual equation:

$$R = -kT \quad - (36)$$

first derived by Einstein uses the assumption that the vacuum is zero. Einstein defined:

$$R := g^{\mu\nu} R_{\mu\nu} \quad - (37)$$

$$T := g^{\mu\nu} T_{\mu\nu} \quad - (38)$$

and used $g^{\mu\nu} g_{\mu\nu} = 4$. $- (39)$

To obtain eq. (36) he started from:

$$R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = k T_{\mu\nu} \quad - (40)$$

and multiplied both sides by $g^{\mu\nu}$ to obtain eq. (36)

Eq. (36) is ECE Theory

In ECE, R is obtained from the Lemma and tetrad potentials:

$$\square q_{\mu}^a = R q_{\mu}^a \quad - (41)$$

where:

$$R := \frac{1}{4} q_{\mu}^a \square q_{\nu}^b - \omega_{\mu b}^a q_{\nu}^b \quad - (42)$$

In ECE: $R := -kT \quad - (43)$

where R is given by eq. (42). Eq. (43) is true in the presence of mass.

7) The most general Noether Theorem is

Therefore:

$$T = -\frac{R}{k} \quad - (44)$$

where R is defined by eq. (42). The
ECE wave equation is ~~the~~:

$$(\square + kT) \psi_{,\mu}^a = 0 \quad - (45)$$

where T is defined by eq. (44).
